

### 372(5): The Complete Set of Equations for the Relativistic Plane Polar Coordinates.

In plane polar coordinates, the velocity is:

$$\underline{v} = \dot{r} \underline{e}_r + r \dot{\phi} \underline{e}_\phi \quad - (1)$$

where  $\underline{e}_r$  and  $\underline{e}_\phi$  are the unit vectors of the plane polar system,  $(r, \phi)$ . Therefore:

$$v^2 = \dot{r}^2 + r^2 \dot{\phi}^2 \quad - (2)$$

The Lorentz factor is:

$$\gamma = \left(1 - \frac{v^2}{c^2}\right)^{-1/2} \quad - (3)$$

and the Lagrangian is:

$$\mathcal{L} = -mc^2 \left(1 - \frac{v^2}{c^2}\right)^{1/2} - U \quad - (4)$$

where  $U$  is the potential energy. For a Coulomb potential:

$$U = -\frac{e^2}{4\pi\epsilon_0 r} \quad - (5)$$

where  $\epsilon_0$  is the vacuum permittivity. The proper Lagrange variables are  $r$  and  $\phi$  and the Euler Lagrange equations are:

$$\frac{\partial \mathcal{L}}{\partial r} = \frac{d}{dt} \left( \frac{\partial \mathcal{L}}{\partial \dot{r}} \right) \quad - (6)$$

$$\frac{\partial \mathcal{L}}{\partial \phi} = \frac{d}{dt} \left( \frac{\partial \mathcal{L}}{\partial \dot{\phi}} \right) \quad - (7)$$

2) The gradient operator is:

$$\underline{\nabla} = \underline{e}_r \frac{\partial}{\partial r} + \underline{e}_\phi \frac{1}{r} \frac{\partial}{\partial \phi} \quad - (8)$$

Quantization occurs via:

$$\text{if } \hat{H} \psi = E \psi \quad - (9)$$

where  $\psi$  is the wavefunction of the relativistic two dimensional problem being considered here. In eq. (9)

$$\hat{H} = \left( \frac{1}{c} \frac{\partial}{\partial t}, -\underline{\nabla} \right) \quad - (10)$$

and

$$\underline{p} = \left( \frac{E}{c}, \underline{p} \right) \quad - (11)$$

Here  $E$  is the relativistic energy:

$$E = \gamma m c^2 \quad - (12)$$

and  $\underline{p}$  is the relativistic momentum:

$$\underline{p} = \gamma m \underline{v} \quad - (13)$$

It follows that:

$$\underline{\nabla} \psi = \frac{\partial \psi}{\partial r} \underline{e}_r + \frac{1}{r} \frac{\partial \psi}{\partial \phi} \underline{e}_\phi \quad - (14)$$

$$\text{if } \frac{\partial \psi}{\partial t} = \gamma m c^2 \psi \quad - (15)$$

$$\text{if } \underline{\nabla} \psi = \gamma m \underline{v} \psi \quad - (16)$$

and

$$\hat{H}^2 \psi = \underline{p}^2 \psi \quad - (17)$$

3) i.e. :

$$\frac{\partial^2 \psi}{\partial r^2} + \frac{1}{r} \frac{\partial \psi}{\partial r} + \frac{1}{r^2} \frac{\partial^2 \psi}{\partial \phi^2} = -\frac{m^2}{\hbar^2} \gamma^2 (\dot{r}^2 + r^2 \dot{\phi}^2) \psi \quad (18)$$

From eq. (16) :

$$-i\hbar \frac{\partial \psi}{\partial r} = \gamma m r \dot{\phi} \psi \quad (19)$$

and

$$-i\hbar \frac{1}{r} \frac{\partial \psi}{\partial \phi} = m r \dot{\phi} \gamma \psi \quad (20)$$

### Computational Scheme

- 1) Find  $r, \phi, \dot{r}$  and  $\dot{\phi}$  from eqs. (6) and (7).
- 2) Find  $\psi(r, \phi)$  from eqs. (19) and (20).
- 3) Find  $\psi(t)$  from eq. (15).
- 4) Check the results using eq. (18).

The Hamiltonian is :

$$H = \gamma m c^2 + U \quad (21)$$

which can be expressed as :

$$H = \frac{p^2}{(1+\gamma)m} + m c^2 - \frac{e^2}{4\pi \epsilon_0 r} \quad (22)$$

Therefore the Hamiltonian operator is :

$$\hat{H}\psi = \left( -\frac{\hbar^2 \nabla^2}{(1+\gamma)m} + m c^2 - \frac{e^2}{4\pi \epsilon_0 r} \right) \psi \quad (23)$$

4) and the energy levels are given by:

$$\hat{H}\psi = E\psi \quad (24)$$

and

$$E = \langle E \rangle = \int \psi^* \hat{H} \psi d\tau \quad (25)$$

These energy levels show two dimensional fine structure.

For the H atom this exercise is repeated using the spherical polar coordinates to give the wave functions of the relativistic H atom and its energy levels showing fine structure. Part will be the subject of the next note.

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