

343(4) : Calculation of the Thomas Phase Shift from the Newtonian Limit.

In previous notes it is shown that the as. t
the Newtonian limit for a rotating frame (r, θ_1) is:

$$r = \frac{d_1}{1 + \epsilon_1 \cos \theta_1} \quad - (1)$$

This produces the non-relativistic orbital velocity:

$$v_{N1}^2 = \frac{L_1^2}{m^2 r^4} \left(r^2 + \left(\frac{dr}{d\theta_1} \right)^2 \right) \quad - (2)$$

The relativistic angular velocity used in Thomas precession theory is:

$$\Omega = \frac{v}{r} \quad - (3)$$

where the relativistic velocity v is defined by:

$$\begin{aligned} v^2 &= \gamma^2 v_{N1}^2 \\ &= \frac{v_{N1}^2}{1 - \frac{v_{N1}^2}{c^2}} \end{aligned} \quad - (4)$$

Therefore the Thomas phase shift is:

$$\Delta \phi = \Omega d\tau - \text{at} \quad - (5)$$

As a UFT 110 :

$$d\tau = \left(1 - \frac{v^2}{c^2} \right)^{1/2} dt \quad - (6)$$

and

2)

$$\Omega = \omega \left(1 - \frac{v_0^2}{c^2} \right)^{-1} - (7)$$

if the infinitesimal line element of ECE2 relativity is rotated. This means that:

$$ds^2 = c^2 d\tau^2 = c^2 dt^2 - dr^2 - r^2 d\theta^2 - (8)$$

is rotated using:

$$d\theta_1 = d\theta + \omega_\theta dt - (9)$$

In the rotating frame, eq. (8) becomes:

$$ds^2 = c^2 d\tau^2 = c^2 dt^2 - dr^2 - r^2 d\theta_1^2 - (10)$$

$$= dt^2 (c^2 - v_{\theta 1}^2)$$

where $v_{\theta 1}^2$ is given by eq. (2).

It follows that the Thomas angular velocity Ω for eqs (3) and (7) must be the same:

$$\Omega = \frac{v}{r} = \omega \left(1 - \frac{v_0^2}{c^2} \right)^{-1} - (11)$$

The Thomas angular velocity is therefore:

$$\Omega = \frac{v_{\theta 1}}{r} - (12)$$

and the Thomas phase shift can be calculated from the orbit (1). The square of the Thomas angular velocity is therefore:

$$\Omega^2 = \frac{1}{r^3} \frac{V_{N1}^2}{1 - V_{N1}^2} \quad - (13)$$

which is calculated from eq. (2) $\frac{V_{N1}^2}{c^2}$

Here:

$$V_{N1}^2 = \frac{L_1^2}{m^2 r^4} \left(r^2 + \left(\frac{dr}{d\theta_1} \right)^2 \right) \quad - (14)$$

where

$$\begin{aligned} \frac{dr}{d\theta_1} &= \frac{d_1 \epsilon_1 \sin \theta_1}{(1 + \epsilon_1 \cos \theta_1)^2} \quad - (15) \\ &= \frac{\epsilon_1 r^2 \sin \theta_1}{d_1} \end{aligned}$$

Here,

$$\theta_1 = \theta + \omega_\theta t \quad - (16)$$

Finally, in order to obtain an idea of the properties of Eq. (1) it can be developed as follows. It can be written as:

$$r = \frac{d_1}{1 + \epsilon_1 \cos(\theta + \omega_\theta t)} \quad - (17)$$

which:

$$\frac{dr}{dt} = \left(\frac{2}{m} (H - U) - \frac{L_1^2}{m^2 r^3} \right)^{1/2} \quad - (18)$$

so that:

$$t = \int \left(\frac{2}{m} (H - U) - \frac{L_1^2}{m^2 r^3} \right)^{-1/2} dr \quad - (19)$$

in which:

$$U = -\frac{mMg}{r} \quad (20)$$

Eqs. (17) and (19) can be solved simultaneously using computer algebra, and a graph of r versus θ plotted. This gives the effect of rotating (r, θ) to give (r, θ_1) or the Newtonian orbit.

From eq. (17):

$$1 + \epsilon_1 \cos(\theta + \omega_0 t) = \frac{\alpha_1}{r} \quad (21)$$

and

$$\cos(\theta + \omega_0 t) = \frac{1}{\epsilon_1} \left(\frac{\alpha_1}{r} - 1 \right) \quad (22)$$

so

$$\theta + \omega_0 t = \cos^{-1} \left(\frac{1}{\epsilon_1} \left(\frac{\alpha_1}{r} - 1 \right) \right) \quad (23)$$

Therefore θ can be expressed in terms of r as follows:

$$\theta = \cos^{-1} \left(\frac{1}{\epsilon_1} \left(\frac{\alpha_1}{r} - 1 \right) \right) - \omega_0 \int \left(\frac{2}{m} (H - U) - \frac{L_1^2}{m^2 r^2} \right)^{-1/2} dr \quad (24)$$

and θ can be plotted against r . This plot can be inverted numerically to give r versus θ . The Newtonian result is:

$$\theta = \cos^{-1} \left(\frac{1}{\epsilon} \left(\frac{\alpha}{r} - 1 \right) \right) \quad (25)$$