

312(1): A Simple Calculation of $m = 8m_0$ from
The Monochromatic Planck Distribution

From eq. (17) of Notes 310 and UFT 310, the flux
density is watts per square metre for a monochromatic
beam is:

$$\Phi = \frac{h\omega}{3c^2\pi^2} \frac{(\omega^2 - \omega_0^2)^{3/2}}{e^y - 1} \quad - (1)$$

where the rest frequency ω_0 is defined by:

$$h\omega_0 = m_0 c^2 \quad - (2)$$

where m_0 is the rest mass of the photon. Using the
lowest observable frequency of

$$f_0 = 0.01 \text{ Hz} \quad - (3)$$

it is found from eq. (2) that:

$$m_0 = 7.39 \times 10^{-53} \text{ kg} \quad - (4)$$

Therefore at visible frequencies ω_0 has an entirely
negligible effect on Φ of eq. (1). So to an excellent
approximation:

$$\Phi = \frac{h\omega^4}{3c^2\pi^2(e^y - 1)} \quad - (5)$$

where

$$y = \frac{h\omega}{kT} \quad - (6)$$

2) Now assume that:

$$\hbar\omega \ll kT \quad - (7)$$

So

$$\Phi \sim \frac{\hbar\omega^4}{3c^2\pi^2 \left(1 + \frac{\hbar\omega}{kT} - 1\right)}$$

$$= \frac{kT\omega^3}{3c^2\pi^2} \quad - (8)$$

Now use the de Broglie equation:

$$E = \hbar\omega = \gamma mc^2 \quad - (9)$$

Then:

$$\Phi = \frac{kT}{\hbar^3 c^2 \pi^2} (\gamma m_0 c^2)^3 \quad - (10)$$

So

$$\Phi = \left(\frac{kT c^4}{3 \hbar^3 \pi^2} \right) (\gamma m_0)^3 \quad - (11)$$

For the sake of argument use:

$$\Phi(x_p) = 1.0 \text{ watt m}^{-2} \quad - (12)$$

Then the moving mass of the photon is:

$$m = \gamma m_0 = \frac{\hbar}{c} \left(\frac{3\pi^2}{c kT} \right)^{1/3} \quad - (13)$$

3) Assume:

$$T = 296 \text{ K}, k = 1.3806 \times 10^{-23} \text{ J K}^{-1}, \quad (14)$$
$$c = 3 \times 10^8 \text{ m s}^{-1}$$

then:

$$m = \gamma m_0 = 1.02 \times 10^{-38} \text{ kg} \quad (15)$$

This means that:

$$\gamma = \frac{1.02 \times 10^{-38}}{7.39 \times 10^{-53}} = 1.38 \times 10^{14} \quad (16)$$

so that v is very close to c because:

$$\gamma = \left(1 - \frac{v^2}{c^2}\right)^{-1/2} \quad (17)$$

This calculation illustrates the fact that the photon mass m_0 produces a speed of light v in the ensemble of free photons that is very close to c . In the writings of the de Broglie/Visser group this is described as $v \approx c$ for all practical purposes (FAPP) for free photons.

4) However when there is absorption of light by an atom or molecule:

$$\frac{\Phi}{\Phi_0} = \exp(-\alpha l) \quad - (18)$$

where the integrated power absorption coefficient is:

$$\alpha = \frac{N}{V} \frac{|\mu_{if}|^2}{6\epsilon_0 \hbar \nu} \quad - (19)$$

Here N is the number of atoms in a volume V , μ_{if} is the transition dipole moment, ϵ_0 the vacuum permittivity, \hbar the reduced Planck constant, l the sample length and ν the velocity of light in a gas of atoms, or a material made up of atoms and molecules.

Then α is $4\pi T/306$:

$$\nu = \nu' + i\nu'' = \frac{c}{n' - i n''} \quad - (20)$$

So:

$$\frac{\nu'}{c} = \frac{n'}{n'^2 + n''^2} \quad - (21)$$

$$\frac{\nu''}{c} = \frac{n''}{n'^2 + n''^2} \quad - (22)$$

and the velocity depends on the dielectric properties of

the material through:

$$n'^2 = \frac{1}{2} (\epsilon' + (\epsilon'^2 + \epsilon''^2)^{1/2}) \quad - (23)$$

$$n''^2 = \frac{4\epsilon''}{(\epsilon' + (\epsilon'^2 + \epsilon''^2)^{1/2})} \quad - (24)$$

The velocity of the Lorentz factor is defined as

$$V_m^2 = |VV^*|^2 = v'^2 + v''^2$$

$$= \frac{c^2}{n'^2 + n''^2} \quad - (25)$$

If the dielectric loss is very small, as in a glass, then:

$$n'^2 \sim \epsilon' \quad - (26)$$

$$n''^2 \sim 0 \quad - (27)$$

and

$$V_m = \frac{c}{n'} \quad - (28)$$

Eq. (28) is the usual equation defining the refractive index:

$$n' = \frac{c}{V_m} \quad - (29)$$

So if:

$$n' = 1.5 \quad - (30)$$

for the sake of argument, then:

$$b) \quad v_m = 1.5 \times 10^8 \text{ ms}^{-1} \quad - (31)$$

In Q, case the moving mass is:

$$\begin{aligned} m &= \left(1 - \frac{1}{4}\right)^{-1/2} m_0 \\ &= \frac{2}{\sqrt{3}} m_0 = 1.155 m_0 \\ &= 8.53 \times 10^{-53} \text{ kg} \end{aligned} \quad - (32)$$

The material has slowed the photon velocity to $1.5 \times 10^8 \text{ ms}^{-1}$ so the moving mass m is very close to the rest mass m_0 .

In order to obtain a fully self-consistent result we:

$$\Phi \sim \left(\frac{kTc^4}{3R^3\pi^2} \right) (\gamma m_0)^3 = \Phi_0 \exp(-dl) \quad - (33)$$

where d is given by eq. (19).

This, will be the subject of the next note
