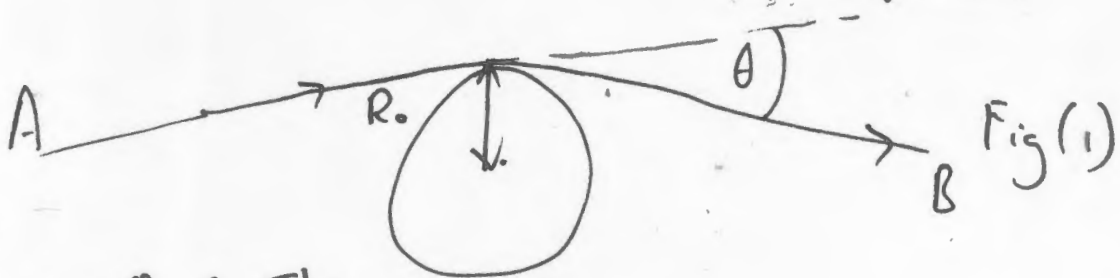


202(6) : Check a Einstein's Starting Equation



Consider the identity:

$$\frac{d\theta}{dr} = \frac{dt}{dr} \quad \text{--- (1)}$$

then
$$d\theta = \left(\frac{d\theta}{dr} \right) dr \quad \text{--- (2)}$$

so
$$\theta = \int \frac{d\theta}{dr} dr \quad \text{--- (3)}$$

When:

$$r = R. \quad \text{--- (4)}$$

the light is at the sun's radius. When:

$$r = \infty \quad \text{--- (5)}$$

the light is far distant from the centre of the sun.

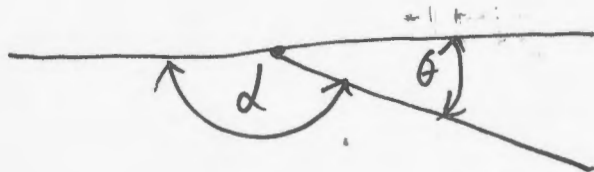
If the light travelled from A to B is a straight line it subtends the angle π (180°).



For a straight line:

$$\frac{d\theta}{dr} = 0 \quad \text{--- (6)}$$

so there is no change in angle.



In the case of light bending:

$$d = \pi - \theta \quad - (7)$$

For a straight line:

$$\frac{d\theta}{dr} = 0 \quad - (8)$$

So $\Delta\theta = \int_{R_0}^{\infty} \frac{d\theta}{dr} dr = 0 \quad - (9)$

This means that:

$$\Delta\theta = \frac{2}{x} \left(\sin^{-1} \frac{1}{\epsilon} - \sin^{-1} \left(\frac{R_0 - d}{\epsilon |R_0|} \right) \right) = 0 \quad - (10)$$

An orbit becomes a straight line when:

$$R_0 \rightarrow \infty \quad - (11)$$

So in eq. (10): $\Delta\theta = 0 \quad - (12)$

self consistently.

From eq. (3):

$$\Delta\theta = \int_{R_0}^{\infty} \left(\frac{d\theta}{dr} \right) dr \quad - (13)$$

and eqs. (11) and (13) give again:

$$\Delta\theta = 0. \quad - (14)$$

3) This is the correct way of calculating the deflection $\Delta\theta$ from the orbit:

$$r = \frac{d}{1 + \epsilon \cos(x\theta)} \quad - (15)$$

The straight line is recovered from this orbit when:

$$r \rightarrow \infty \quad - (16)$$

which means that the light is infinitely away from the sun.

This means that there is another problem in the Einstein calculation.

In Wald's notation the deflection $\Delta\theta$ is given by eq. (6.3.39), page 145 of R.M. Wald, "General Relativity" (Chicago, 1984):

$$\Delta\theta = 2 \int_0^{R_0} \frac{du}{(b^2 - u^2 + 2Mu^3)^{1/2}} \quad - (17)$$

and the straight line is recovered with:

$$M = ? 0, \quad R_0 = ? b \quad - (18)$$

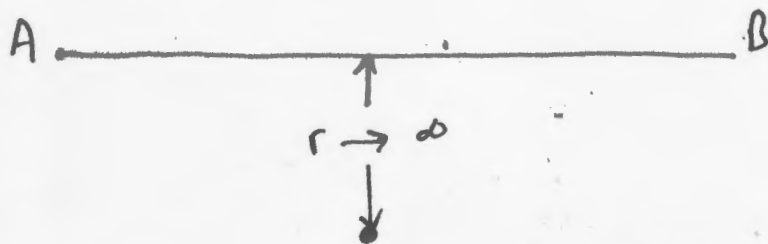
$$\text{to give } \Delta\theta = \pi. \quad - (19)$$

$$\text{This means that: } \frac{dr}{d\theta} \neq ? 0 \quad - (20)$$

for a straight line!

This is under absurd result.

4) In case of no deflection:



In this case: $\frac{dr}{d\theta} = \left(\frac{\epsilon x}{L}\right) r^2 \sin(x\theta) \quad - (21)$

So $\frac{d\theta}{dr} \xrightarrow{r \rightarrow \infty} 0 \quad - (22)$

because $\frac{d\theta}{dr} = \left(\frac{d}{\epsilon x}\right) \frac{1}{r^2} \cdot \frac{1}{\sin(x\theta)} \quad - (23)$

See it another way:

$$\Delta\theta = \int_{-\infty}^{\infty} \left(\frac{d\theta}{dr}\right) dr = 0 \quad - (24)$$

The result (22) is understandable from the fact that it is a straight line there is no dependence of r or angle.

So the complete expression for deflection is:

$$\Delta\theta = \frac{2}{x} \left(\sin^{-1} \frac{1}{\epsilon} - \sin^{-1} \left(\frac{1}{\epsilon} \left(1 - \frac{d}{R_0} \right) \right) \right) \quad - (25)$$

