

1)

CALCULATION OF THE EDDINGTON EFFECT IN GENERAL RELATIVITY AND NEWTONIAN DYNAMICS.

1) In Newtonian dynamics consider a mass m subjected to the gravitational pull of the sun, of mass M . The force between these two masses is:

$$\underline{F} = \frac{-mM\Gamma}{r^2} \underline{k} \quad - (1)$$

Integrate eqn (1) to obtain:

$$T = \frac{1}{2} m v^2 = \frac{mM\Gamma}{r} \quad - (2)$$

This gives:

$$v^2 = \frac{2M\Gamma}{r} \quad - (3)$$

and:

$$\begin{aligned} c^2 - v^2 &= (c-v)(c+v) \\ &= c^2 \left(1 - \frac{2M\Gamma}{rc^2} \right) \end{aligned} \quad - (4)$$

If $v \ll c$, then:

$$c-v \sim c \left(1 - \frac{2M\Gamma}{rc^2} \right)^{1/2}$$

$$c-v \sim c \left(1 - \frac{2M\Gamma}{rc^2} \right) \quad - (5)$$

2) In general relativity the same result is obtained from the Schwarzschild metric:

$$2) \quad ds^2 = \left(1 - \frac{2GM}{rc^2}\right) c^2 dt^2 - \left(1 - \frac{2GM}{rc^2}\right)^{-1} \\ - r^2 d\theta^2 - r^2 \sin^2\theta d\phi^2. \quad - (6)$$

For any mass m the velocity v in the frame of the observer is:

$$v_0 = \frac{dr}{dt} = \frac{\left(1 - \frac{2GM}{rc^2}\right)^{1/2}}{\left(1 - \frac{2GM}{rc^2}\right)^{-1/2}} \frac{dr'}{dt'} \quad - (7)$$

$$v_0 = \left(1 - \frac{2GM}{rc^2}\right) v' \quad - (8)$$

For a photon its speed appears to be slowed by

$$\boxed{v_{0.6} = \left(1 - \frac{2GM}{rc^2}\right) c} \quad - (9)$$

Eqn (9) is observed rather than eqn. (5).

To interpret eqn. (7) recall that the length contraction in special relativity is:

$$l = \gamma l' \quad - (10)$$

and that the relativistic momentum is:

$$p = \gamma p' \quad - (11)$$

The time dilation is:

3)

$$t = \frac{1}{\gamma} t' \quad - (12)$$

The proper time is:

$$\tau = t' \quad - (13)$$

is ~~the~~ ^{moving} ~~observed~~ frame, so:

$$ds^2 = c^2 d\tau^2 = c^2 dt^2 - dx^2 - dy^2 - dz^2 \quad - (14)$$

is the Minkowski metric.

The Schwarzschild metric is ~~the~~ a generalization of the Minkowski metric. Eqn. (7) follows from ~~this~~ this.